

# Magnetic moments of light nuclei

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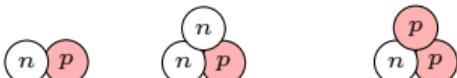
Summary

# Introduction

# Current status of few-nucleon magnetic form factors

Magnetic moments of light nuclei

## Experiment



- ▶ Magnetic moments of deuteron, triton, and helion known extremely precisely  
e.g.  $\mu_d = 0.8574382346(53)\mu_N$  [Puchalski\\_2015](#)
- ▶ Magnetic radii known poorly  
(muonic spectroscopy measurements discussed [Antognini:2015vxo](#))
- ▶ Some data of deuteron magnetic form factor, less for 3N ones

## Chiral EFT

- ▶ Other recent studies [Marcucci:2015rca](#); [Schiavilla:2018udt](#); [Seutin:2021hls](#) are to N<sup>3</sup>LO and lack proper error analysis or put different focus
- ▶ Can use the same setup as for charge form factors [Arseniy's talk](#), in particular:
  - N<sup>4</sup>LO<sup>+</sup> chiral 2N forces
  - N<sup>2</sup>LO chiral 3N forces (plus selected N<sup>4</sup>LO 3N forces) [Patrick's talk](#)
  - uncertainty estimation (most importantly truncation errors)
- ▶ Spatial components of isoscalar vector current operators also known up through N<sup>4</sup>LO

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- ▶ All ingredients to perform **precision calculations** of magnetic observables
- ▶ Accurate experimental magnetic moments provide **benchmarks** for ChEFT  
→ for **light as well as heavier nuclei**
- ▶ **Improve predictions** of magnetic form factors and radii
- ▶ Fix unknown LECs for reuse in other calculations

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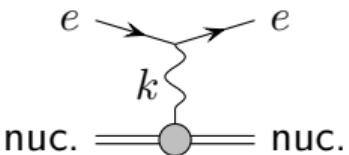
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# Magnetic form factors

Scattering cross section:  $\frac{d\sigma}{d\Omega} = \left(\frac{d\sigma}{d\Omega}\right)_{\text{NS}} F(Q^2, \theta)^2$

$$Q^2 := -k_\mu k^\mu$$



Spin-1 particle (deuteron)

$$\begin{array}{c} F \\ \swarrow \quad \downarrow \quad \searrow \\ G_C \quad G_Q \quad G_M \end{array}$$

Spin- $\frac{1}{2}$  particle (triton, helion)

$$\begin{array}{c} F \\ \swarrow \quad \searrow \\ F_C \quad F_M \end{array}$$

$$\begin{aligned} F_M^S &= \frac{1}{2} (F_M^{^3\text{He}} + F_M^{^3\text{H}}) \\ F_M^V &= \frac{1}{2} (F_M^{^3\text{He}} - F_M^{^3\text{H}}) \end{aligned}$$

Magnetic moments:  $\mu \propto X_M(0)$

Magnetic radii:  $r_M^2 \propto X'_M(0)$

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# Calculation

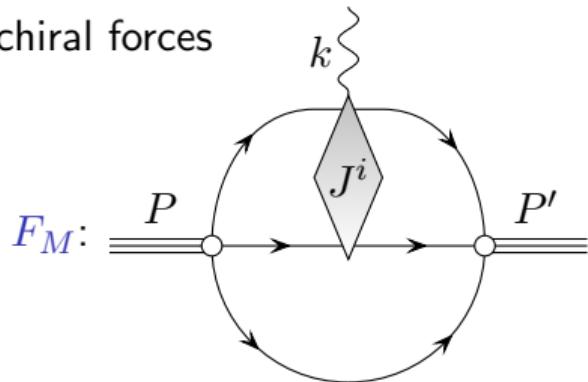
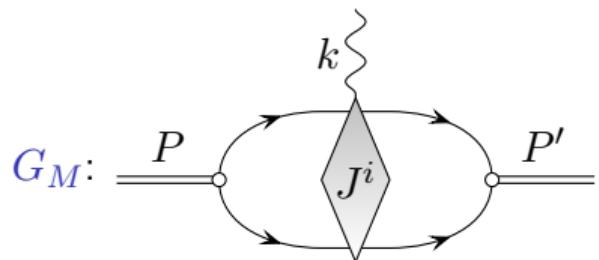
# Calculation of magnetic form factors

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Convolutions  $\langle P', \lambda' | J^i | P, \lambda \rangle$  of

a) spatial components of the vector current operator  $J^i$

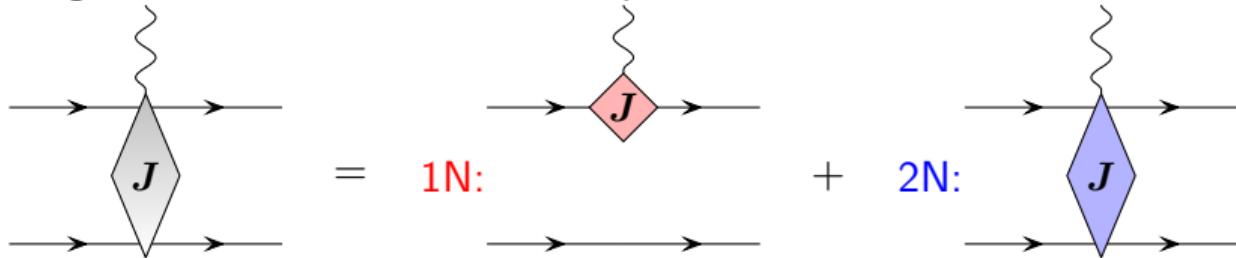
b) with wave functions  $|P, \lambda\rangle$  obtained from SMS chiral forces



1. Solution of Schrödinger and Faddeev equations in PW basis
2. Decomposition of current operators in PW basis
3. Analytic integration over angles (either directly or by Fourier transform to coordinate space)
4. Cheap numeric integration over absolute values of momenta or coordinates
5. Fitting of unknown LECs
6. Error analysis

# Current operators

- ▶ Single-nucleon and two-nucleon operators:

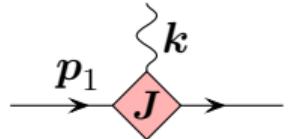


3N operators appear at higher orders

- ▶ No leading-order (LO,  $q^{-3}$ ) three-current operator in our power counting, first contribution at NLO ( $q^{-1}$ )
- ▶ Operators at  $N^3LO$  only known in dim-reg [Kolling:2011mt](#); [Krebs:2019aka](#)
- ▶ Current operators require **consistent regularization** with potential
  - for isovector currents beyond  $N^2LO$  still work in progress [Hermann's talk](#)
  - ⇒ **Focus on isoscalar currents and observables**

# Single-nucleon contributions

From NLO:  $J_{1N} = -\frac{ie}{2m_N} \mathbf{k} \times \boldsymbol{\sigma} \mathcal{G}_M + \frac{e}{2m_N} (2\mathbf{p}_1 + \mathbf{k}) \mathcal{G}_E$



## Single-nucleon Sachs FFs $\mathcal{G}_E$ and $\mathcal{G}_M$

- Keep unexpanded and replace by phenomenological parametrization:

[Ye:2017gyb](#), check consistency with [Belushkin:2006qa](#); [Lin:2021umz](#)

- $\mathcal{G}_{E(M)} = \frac{1+\tau^3}{2} \mathcal{G}_{E(M)}^p + \frac{1-\tau^3}{2} \mathcal{G}_{E(M)}^n$

- $\mathcal{G}_{E(M)}^S := \mathcal{G}_{E(M)}^p + \mathcal{G}_{E(M)}^n$

$$\mathcal{G}_{E(M)}^V := \mathcal{G}_{E(M)}^p - \mathcal{G}_{E(M)}^n$$

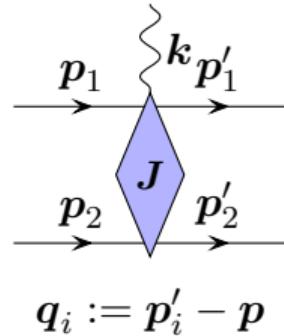
No further 1N contributions up to and including N<sup>4</sup>LO  
(also no relativistic boost corrections to rest-frame wave functions up to this order)

# Two-nucleon contributions

From NLO:

$$\mathbf{J}_{1\pi}^{\text{NLO}} = ie\mathcal{G}_E^V \frac{g_A^2}{4F_\pi^2} [\boldsymbol{\tau}_1 \times \boldsymbol{\tau}_2]^3 \frac{\boldsymbol{\sigma}_2 \cdot \mathbf{q}_2}{\mathbf{q}_2^2 + M_\pi^2} \left( \mathbf{q}_1 \frac{\boldsymbol{\sigma}_1 \cdot \mathbf{q}_1}{\mathbf{q}_1^2 + M_\pi^2} - \boldsymbol{\sigma}_1 \right)$$

- ▶ parameter-free
- ▶ fully **isovector**



From N<sup>3</sup>LO:  $\mathbf{J}_{1\pi}^S = ie2\mathcal{G}_E^S \frac{-3g_A}{F_\pi^2} d_9 \frac{\boldsymbol{\sigma}_2 \cdot \mathbf{q}_2}{\mathbf{q}_2^2 + M_\pi^2} \mathbf{k} \times \mathbf{q}_2 \rightarrow \text{isoscalar } 1\pi \text{ exch.}$

$$\mathbf{J}_{\text{cont}}^S = ie\mathcal{G}_E^S L_2 (\boldsymbol{\sigma}_1 + \boldsymbol{\sigma}_2) \times \mathbf{k} \rightarrow \text{isoscalar contact}$$

$$\mathbf{J}_{\text{cont}}^V = ie\mathcal{G}_E^V L_1 \tau_1^3 (\boldsymbol{\sigma}_1 - \boldsymbol{\sigma}_2) \times \mathbf{k} \rightarrow \text{single isovector contact}$$

- ▶  $d_9$  is determined in  $\pi N$  ChPT
- ▶  $L_2, L_1$  are determined in this study

No further **isoscalar** contributions at N<sup>4</sup>LO

# Regularization of two-nucleon contributions

Regularization  $1\pi$  exchange contributions:

Replace propagators the same way as in the SMS potential

$$\frac{1}{q^2 + M_\pi^2} \rightarrow \frac{e^{-\frac{q^2 + M_\pi^2}{\Lambda^2}}}{q^2 + M_\pi^2}, \quad \frac{1}{(q^2 + M_\pi^2)^2} \rightarrow \frac{e^{-\frac{q^2 + M_\pi^2}{\Lambda^2}}}{(q^2 + M_\pi^2)^2} + \frac{1}{\Lambda^2} \frac{e^{-\frac{q^2 + M_\pi^2}{\Lambda^2}}}{q^2 + M_\pi^2}$$

Regularization of contact contributions:

Transversal  $\mathbf{k} \cdot \mathbf{J} = 0 \Rightarrow$  current conservation does not help

Comparing with similar terms at higher order or other isovector terms at the same order, the most consistent approach appears to be

$$\begin{aligned} \mathbf{J}_{\text{cont}} \rightarrow \frac{1}{4} & \left( \exp \left[ -\frac{(\mathbf{p} + \mathbf{k}/2)^2 + \mathbf{p}'^2}{\Lambda^2} \right] + \exp \left[ -\frac{(\mathbf{p} - \mathbf{k}/2)^2 + \mathbf{p}'^2}{\Lambda^2} \right] \right. \\ & \left. + \exp \left[ -\frac{\mathbf{p}^2 + (\mathbf{p}' + \mathbf{k}/2)^2}{\Lambda^2} \right] + \exp \left[ -\frac{\mathbf{p}^2 + (\mathbf{p}' - \mathbf{k}/2)^2}{\Lambda^2} \right] \right) \mathbf{J}_{\text{cont}}, \end{aligned}$$

with  $\mathbf{p}^{(l)} = (\mathbf{p}_1^{(l)} - \mathbf{p}_2^{(l)})/2$

→ to be verified once results from higher-derivative regularization are available

# Fits and uncertainties

Fitting of unknown LECs from N<sup>3</sup>LO operators:

- ▶ Isoscalar 1 $\pi$ -exch. LEC  $d_9 \xrightarrow{\text{set to}} 0$   
compatible with determination in  $\pi N$  sector [Rijneveen:2021bfw](#)
- ▶ Isoscalar contact LEC  $L_2 \xrightarrow{\text{fit to}} \mu_d^{\text{exp}}$
- ▶ Isovector contact LEC  $L_1 \xrightarrow{\text{fit to}} \mu_3^{V,\text{exp}}$

Error analysis:

- ▶ Bayesian analysis considering chiral expansion order by order  
[Furnstahl:2015rha](#); [Melendez:2017phj](#); [Wesolowski:2018ljj](#); [Epelbaum:2019zqc](#)
- ▶ No LO for magnetic observables  $\rightarrow$  modified  $\bar{C}_{0.5-10}^{650}$  model starting at NLO:  
$$X_M = X_M^{(2)} + \Delta X_M^{(3)} + \Delta X_M^{(4)} + \dots =: X_M^{\text{ref}} (c_2 q^2 + c_3 q^3 + c_4 q^4 + \dots)$$
with  $X_M^{\text{ref}} = \max \left( \frac{|X_M^{(2)}|}{q^2}, \frac{|\Delta X_M^{(3)}|}{q^3}, \frac{|\Delta X_M^{(4)}|}{q^4} \right)$
- ▶ Inclusion of uncertainties beyond truncation errors is work in progress  
(assumed to be less dominant)

## Results (all preliminary)

# Deuteron magnetic form factor

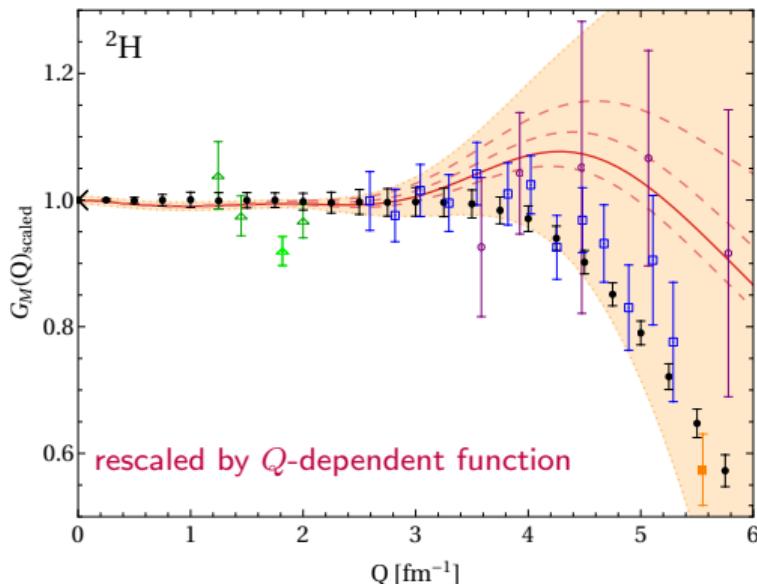
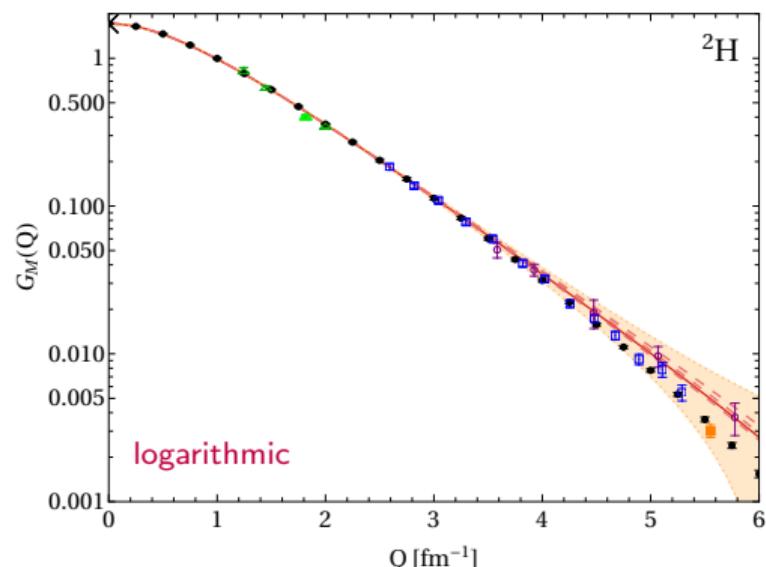


Magnetic moments of light nuclei

Prediction of FF shape at N<sup>4</sup>LO with  $\Lambda = 450$  MeV

band is 68% DoB truncation error, dashed are other cutoffs

Experimental data (colored) and parametrization by Sick (black) from [Marcucci:2015rca](#)



→ Excellent description of data within truncation error

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# 3N isoscalar magnetic form factor



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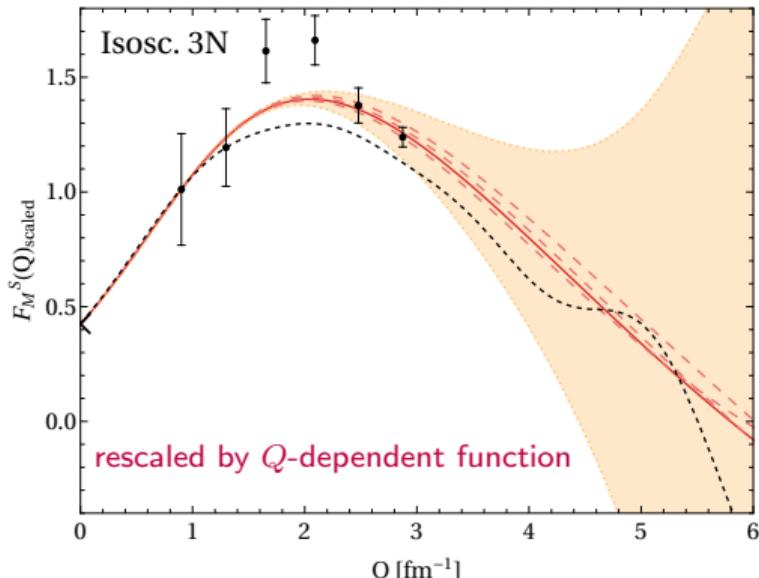
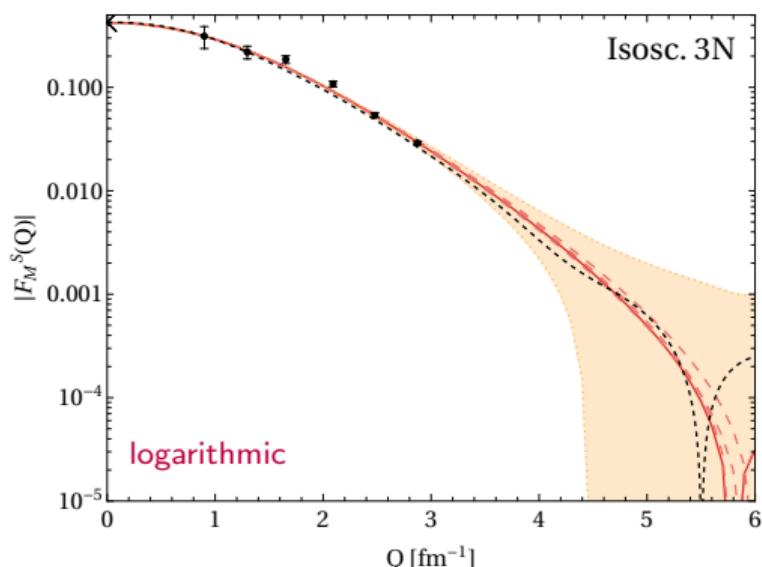
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Prediction of FF shape and magnetic moment ( $N^4LO$ ,  $\Lambda = 450$  MeV)

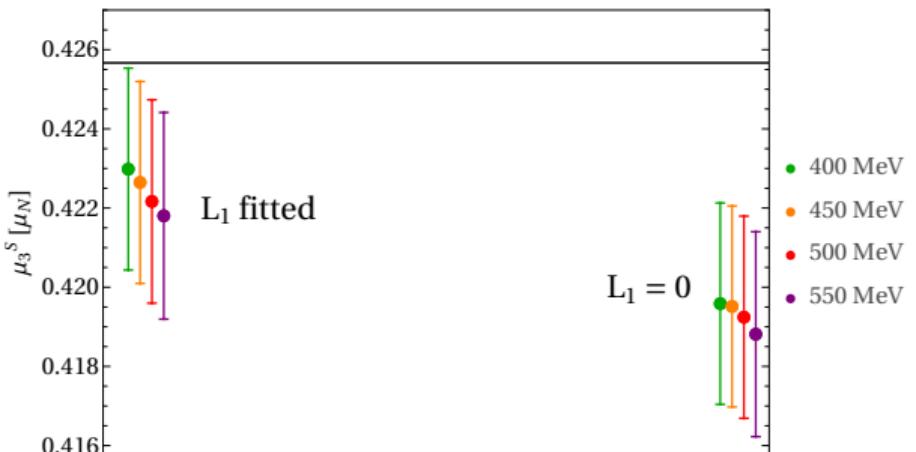
Data from [Beck:1987zz](#) and parametrization by Sick (black dashed) from [Marcucci:2015rca](#)



→ Prediction overall consistent with (poor) data

# Isoscalar magnetic moment

Prediction of isoscalar 3N magnetic moment at N<sup>4</sup>LO, 68% DoB truncation errors



- Successful precision test of ChEFT (regarding order and DoB)
- Essential to include isovector current at N $\geq$ 3LO, fully covered by  $L_1$  here

# Isoscalar magnetic moment and magnetic radii

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Results for magnetic moment and radius predictions and corresponding LECs  
(N<sup>4</sup>LO, 68% DoB truncation errors)

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$\Lambda$ in MeV	$L_2$ in $F_\pi^{-2} \Lambda_b^{-2}$	$L_1$ in $F_\pi^{-2} \Lambda_b^{-2}$	$\mu_3^S$ in $\mu_N$	$r_{Md}^2$ in fm <sup>2</sup>	$(r_{M3}^S)^2$ in fm <sup>2</sup>
400	0.023	1.285	0.4230(26)	4.481(27)	2.283(14)
450	0.052	1.201	0.4226(26)	4.481(27)	2.279(14)
500	0.079	1.187	0.4222(26)	4.481(27)	2.273(14)
550	0.109	1.225	0.4218(26)	4.480(27)	2.261(15)
experiment	—	—	0.425 668 622(6)	4.29(7)	2.1(24)

- ▶ (Isovector)  $L_1$  large compared to (isoscalar)  $L_2$  but of natural size
- ▶ Predicted radii much more precise than current experimental knowledge

# Isovector and individual 3N FFs &



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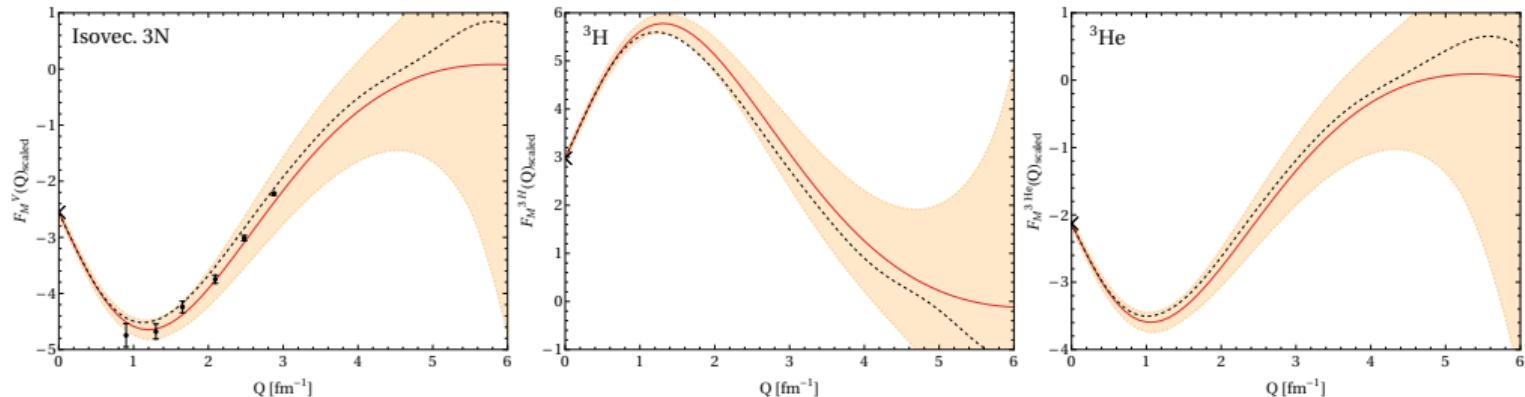
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Prediction of FF shapes and  $^3\text{H}$  and  $^3\text{He}$  magnetic moments at (incomplete)  $\text{N}^3\text{LO}$



$\Lambda$ in MeV	$\mu_{^3\text{H}}$ in $\mu_N$	$\mu_{^3\text{He}}$ in $\mu_N$	$(r_{M3}^V)^2$ in $\text{fm}^2$	$r_{^3\text{H}}^2$ in $\text{fm}^2$	$r_{^3\text{He}}^2$ in $\text{fm}^2$
450	2.98(7)	-2.13(6)	3.53(8)	3.35(8)	3.78(9)
500	2.98(7)	-2.13(6)	3.54(8)	3.36(8)	3.79(9)
experiment	2.978 962 471(10)	-2.127 625 227(8)	3.6(4)	3.4(7)	3.90(19)

→ Good agreement with data, including just a single  $\text{N}^3\text{LO}$  isovector contribution

# Summary

- ▶ Calculation of *isoscalar* magnetic few-nucleon observables pushed to  $N^4\text{LO}$
- ▶ Calculation of *isovector* magnetic few-nucleon observables pushed beyond  $N^2\text{LO}$
- ▶ Good agreement with experimental data, in particular
  - successfully benchmarked ChEFT with isoscalar 3N magnetic moment
  - unprecedented precision for magnetic radii

## Outlook

- ▶ Room for improvements, in particular for isovector and individual  ${}^3\text{H}$ ,  ${}^3\text{He}$  form factors via consistently regularized (isovector)  $N \geq 3\text{LO}$  current operators
- ▶ Calculation should be easily extensible to magnetic properties of larger nuclei,  
**magnetic moments can provide ChEFT benchmark in addition to charge radii**  
→ We can provide “magnetic moment operators” to be directly plugged into existing frameworks.

# Supplementary slides

# Fitting $d_9$ to $\mu_3^S$ at N<sup>4</sup>LO

- ▶ Having three precise experimental values  $\mu_d$ ,  $\mu_3^S$ ,  $\mu_3^V$ , one can principally fit all LECs  $L_2$ ,  $L_1$ , and  $d_9$  simultaneously as done in Schiavilla:2018udt.
- ▶ Similar as there, we obtain a range of

$$d_9 = 0.3 \dots 0.5 \text{ GeV}^{-2}$$

by varying cutoffs.

( $d_9$  actually cutoff-independent, since  $\pi N$  quantity)

- ▶ Quite sizable contribution but practically completely cancelled by  $L_2$  contribution
  - sum of both is very small and form factors as well as radii change only insignificantly compared to  $d_9 = 0$
- ▶ Change to  $\mu_3^S$  is only within truncation error

⇒ Unable to reliably extract non-zero  $d_9$  from few-nucleon data

# $N^5LO$ contact current

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Supplementary slides

Fitting  $d_9$

Beyond  $N^4LO$

Isovector observables

1N FF parametrizations

3N forces

Magnetic moments order by order

Contributions to magnetic moments

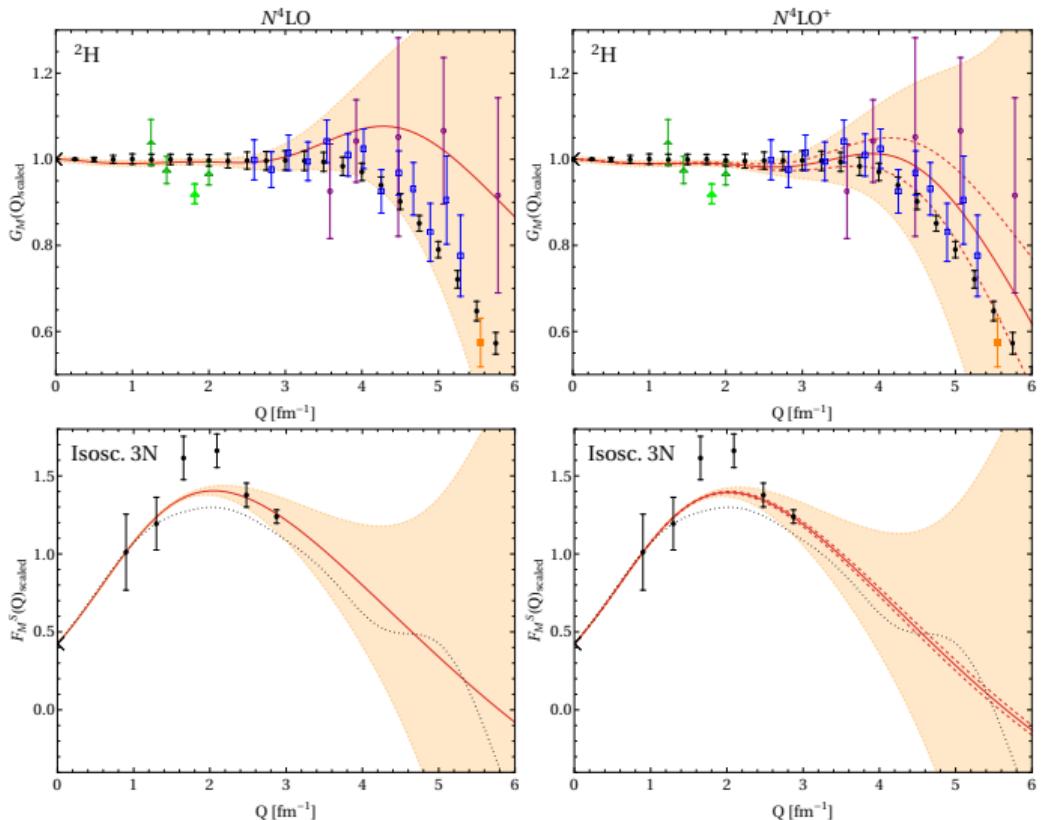
References

- Derivable analogously to higher-order *charge* corrections in [Filin:2020tcs](#) via unitary transformation  $U = e^{AT_1+BT_2+CT_3}$  of the leading current operator:

$$\hat{\mathbf{J}}_{\text{cont}}^{\text{N}^5\text{LO}} = \hat{U}^\dagger \hat{\mathbf{J}}_{1\text{N}} \hat{U} - \hat{\mathbf{J}}_{1\text{N}} \simeq [\hat{\mathbf{J}}_{1\text{N}}, \hat{\mathcal{T}}]$$

- Regularization of the generators  $T_i$  consistent with the chiral potential from [Reinert:2017usi](#) as  $T_i \rightarrow T_i \exp\left(-\frac{\mathbf{p}^2 + \mathbf{p}'^2}{\Lambda^2}\right)$
- Gives three more LECs, only two contribute to deuteron
- Fitting to  $G_M$  data yields high correlation of these LECs (not enough data from isoscalar 3N form factor)
- Fit only one of them to  $G_M$  shape (no contribution to magnetic moments)
- $\chi^2$ -fit including experimental and theoretical errors (1N form factor parametrization and truncation [iteratively])

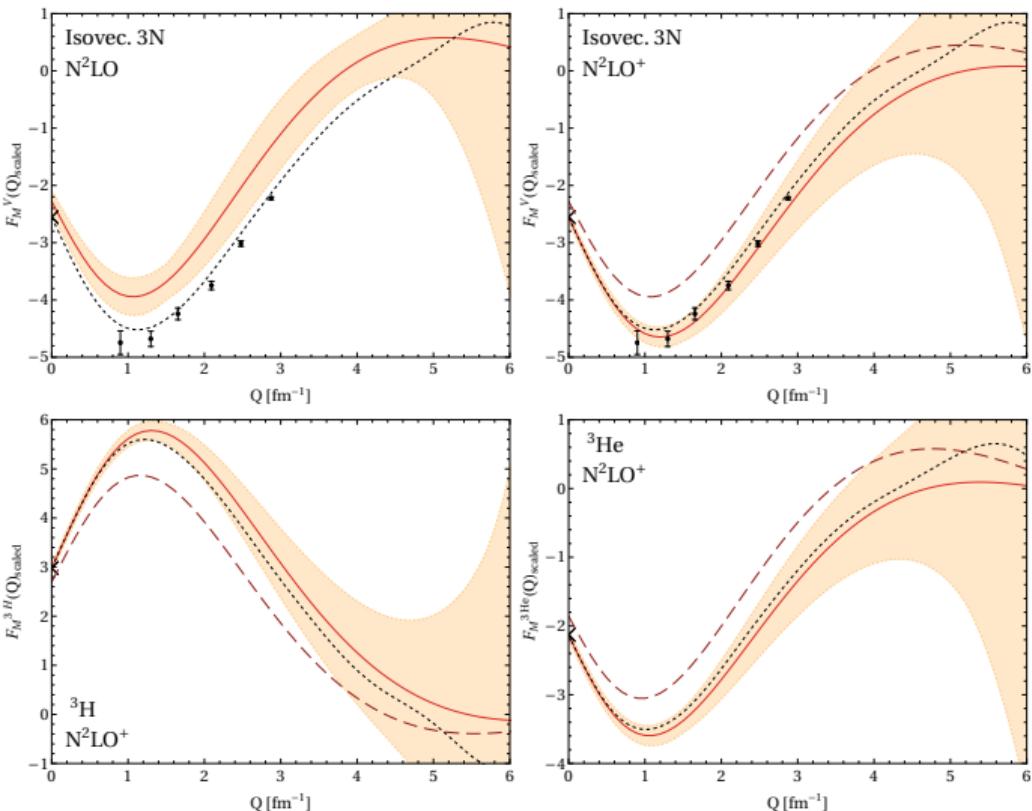
# Results beyond N<sup>4</sup>LO



- ▶ Slightly improved description of  $G_M$  data after fitting single LEC
- ▶ Statistical uncertainty (red dashed) within truncation error
- ▶ Effect on  $F_M^S$  negligible
- ▶ Radii only get a few per mille larger
- ⇒ N<sup>4</sup>LO prediction already does a good job

# Details on isovector and individual 3N FFs

- ▶ Up through N<sup>2</sup>LO all operators included
- ▶ At N<sup>3</sup>LO only single isovector operator included in addition
- ▶ Fitting  $L_1$  seems to do a good job compared to  $L_1 = 0$  (dashed)
- ▶ Magnetic moment and radius predictions well in line with data but large (few-percent) truncation errors



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Beyond N<sup>4</sup>LO

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3N forces

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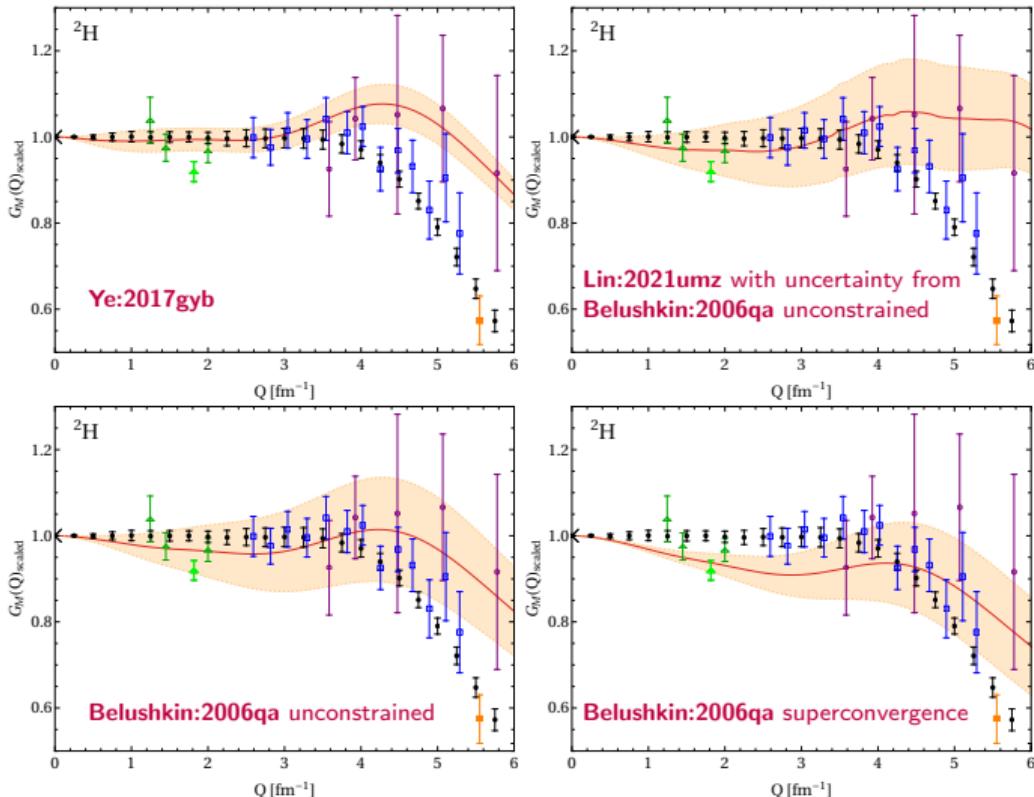
References

# Effects of changing 1N FF parametrizations

Scaled  $G_M$  for different parametrizations of the 1N Sachs form factors  $\mathcal{G}_E, \mathcal{G}_M$

(bands are propagated uncertainties of the parametrizations)

⇒ Within errors not relevant if Ye:2017gyb or Lin:2021umz are used though deviation is larger than for deuteron charge form factors (due to  $\mathcal{G}_M$  being more important here)



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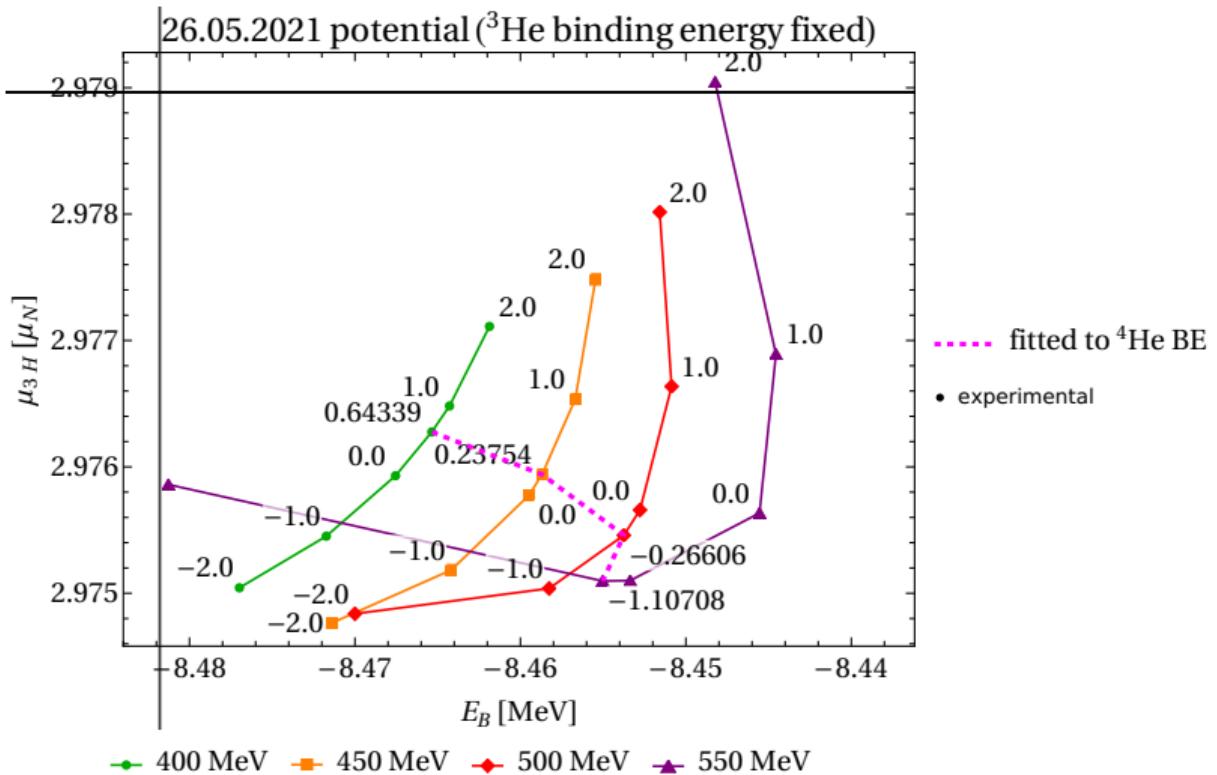
Magnetic moments order by order

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# Correlation of $^3\text{H}$ binding energy and magnetic moment

Varying  $c_{E_1}$  3N force contribution yields similar correlation as for charge radii:



# Chiral convergence of magnetic moments

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$$\mu_d \ 0.857\,438\,234\,6(53)$$

$\Lambda$	NLO	$N^2\text{LO}$	$N^3\text{LO fit}$	$N^4\text{LO fit}$
400	0.86(31)	0.86(7)	0.857(18)	0.857(5)
450	0.86(31)	0.85(7)	0.857(18)	0.857(5)
500	0.86(31)	0.85(7)	0.857(18)	0.857(5)
550	0.85(31)	0.84(7)	0.857(18)	0.857(5)

$$\mu_3^S \ 0.425\,668\,622(6)$$

$\Lambda$	NLO	$N^2\text{LO}$	$N^3\text{LO}$	$N^4\text{LO}$
400	0.42(15)	0.416(34)	0.423(9)	0.4230(26)
450	0.42(15)	0.412(33)	0.423(9)	0.4226(26)
500	0.42(15)	0.407(33)	0.424(9)	0.4222(26)
550	0.42(15)	0.403(33)	0.424(9)	0.4218(26)

$$\mu_3^V \ -2.553\,293\,849(6)$$

$\Lambda$	NLO	$N^2\text{LO}$	$N^3\text{LO fit}$	$N^4\text{LO fit}$
400	-2.3(8)	-2.26(18)	-2.55(7)	-2.553(20)
450	-2.3(8)	-2.28(18)	-2.55(7)	-2.553(19)
500	-2.3(8)	-2.30(18)	-2.55(6)	-2.553(18)
550	-2.3(9)	-2.32(19)	-2.55(6)	-2.553(17)

$$\mu_{^3\text{H}} \ 2.978\,962\,471(10) \text{ beginning from } N^3\text{LO incomplete}$$

$\Lambda$	NLO	$N^2\text{LO}$	$N^3\text{LO}$	$N^4\text{LO}$
400	2.7(10)	2.67(21)	2.98(8)	2.976(21)
450	2.7(10)	2.69(22)	2.98(7)	2.976(20)
500	2.7(10)	2.70(22)	2.98(7)	2.975(19)
550	2.8(10)	2.72(22)	2.98(7)	2.975(19)

$$\mu_{^3\text{He}} \ -2.127\,625\,227(8) \text{ beginning from } N^3\text{LO incomplete}$$

$\Lambda$	NLO	$N^2\text{LO}$	$N^3\text{LO}$	$N^4\text{LO}$
400	-1.8(7)	-1.84(15)	-2.13(7)	-2.130(19)
450	-1.9(7)	-1.86(15)	-2.13(6)	-2.131(18)
500	-1.9(7)	-1.89(15)	-2.13(6)	-2.131(16)
550	-1.9(7)	-1.91(15)	-2.13(5)	-2.131(15)

# Individual contributions to magnetic moments

Percentages of individual contributions to deuteron and isoscalar 3N magnetic moments relative to the experimental value at N<sup>4</sup>LO

$\mu_d$ in %:	$\Lambda$	1N contrib.	$L_2$ -term	truncation error		
400	99.72	0.28	0.60			
450	99.36	0.64	0.60			
500	99.03	0.97	0.60			
550	98.74	1.26	0.60			

$\mu_3^S$ in %:	$\Lambda$	1N contrib.	NLO OPE	$L_2$ -term	$L_1$ -term	missing	truncation error
400	97.91	-0.07	0.72	0.80	0.63	0.60	
450	96.90	0.02	1.63	0.74	0.71	0.60	
500	95.96	0.10	2.44	0.69	0.82	0.60	
550	95.10	0.15	3.14	0.70	0.91	0.61	

→ Similar as for charge radii, 2N contributions drastically reduce cutoff dependence and their importance grows with number of nuclei

# References

Magnetic  
moments of  
light nuclei

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Supplementary  
slides

References

- A. Antognini (Dec. 2015).
- D. Beck *et al.*, *Phys. Rev. Lett.* **59**, 1537–1540 (1987).
- M. A. Belushkin, H.-W. Hammer, U.-G. Meißner, *Phys. Rev. C* **75**, 035202 (2007).
- S. D. Drell, J. D. Walecka, *Annals Phys.* **28**, 18–33 (1964).
- E. Epelbaum *et al.*, *Eur. Phys. J. A* **56**, 92 (2020).
- A. A. Filin, D. Möller, V. Baru, E. Epelbaum, H. Krebs, P. Reinert, *Phys. Rev. C* **103**, 024313 (2021).
- R. J. Furnstahl, N. Klco, D. R. Phillips, S. Wesolowski, *Phys. Rev. C* **92**, 024005 (2015).
- M. Garcon, J. W. Van Orden, *Adv. Nucl. Phys.* **26**, 293 (2001).
- R. A. Gilman, F. Gross, *J. Phys. G* **28**, R37–R116 (2002).
- S. Kölling, E. Epelbaum, H. Krebs, U.-G. Meißner, *Phys. Rev. C* **84**, 054008 (2011).
- H. Krebs, E. Epelbaum, U.-G. Meißner, *Few Body Syst.* **60**, 31 (2019).
- Y.-H. Lin, H.-W. Hammer, U.-G. Meißner, *Eur. Phys. J. A* **57**, 255 (2021).
- L. E. Marcucci *et al.*, *J. Phys. G* **43**, 023002 (2016).
- J. A. Melendez, S. Wesolowski, R. J. Furnstahl, *Phys. Rev. C* **96**, 024003 (2017).
- Y. I. Neronov, N. N. Seregin, *J. Exp. Theor. Phys.* **115**, 777–781 (Nov. 2012).
- M. Puchalski, J. Komasa, K. Pachucki, *Phys. Rev. A* **92** (Aug. 2015).
- P. Reinert, H. Krebs, E. Epelbaum, *Eur. Phys. J. A* **54**, 86 (2018).
- N. Rijnveen, A. M. Gasparyan, H. Krebs, E. Epelbaum (Aug. 2021).
- R. Schiavilla *et al.*, *Phys. Rev. C* **99**, 034005 (2019).
- R. Seutin, PhD thesis, Tech. U., Dortmund (main), Tech. U., Dortmund (main), 2021.
- S. Wesolowski, R. J. Furnstahl, J. A. Melendez, D. R. Phillips, *J. Phys. G* **46**, 045102 (2019).
- Z. Ye, J. Arrington, R. J. Hill, G. Lee, *Phys. Lett. B* **777**, 8–15 (2018).
- D. R. Yennie, M. M. Lévy, D. G. Ravenhall, *Reviews of Modern Physics* **29**, 144–157 (Jan. 1957).