Hydrodynamics — A tool for Strongly Coupled Systems

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# Hydrodynamics

# Hydrodynamics in a Nutshell

Hydrodynamics = Conservation of Energy and Momentum

$$\partial_{\mu}T^{\mu\nu}=0$$

T<sup>μν</sup> = Energy Momentum Tensor. For fluids, can only depend on

 $\epsilon(P, T), u^{\mu}, g^{\mu\nu}$ , gradients thereof

• In equilibrium and local rest frame  $(u^{\mu} = (1, 0, 0, 0))$ ,

$$T_{\rm eq,LRF}^{\mu\nu} = {\rm diag}(\epsilon, P, P, P)$$

#### Hydrodynamics in a Nutshell

Only possible form

$$T^{\mu\nu} = (\epsilon + P) u^{\mu} u^{\nu} - P g^{\mu\nu} + \text{gradients}$$

Gradients:

gradients = 
$$\eta \nabla^{<\mu} u^{\nu>} + \zeta \left( g^{\mu\nu} - u^{\mu} u^{\nu} \right)$$

- η... Shear viscosity coefficient
- ζ... Bulk viscosity coefficients (usually neglected)

#### Hydrodynamics in a Nutshell

• Neglect  $\eta, \zeta$  and take non-relativistic limit:

$$u_{\nu}\partial_{\mu}T^{\mu\nu} = \mathbf{0} \to \partial_{t}\rho + \partial_{i}\left(\rho\mathbf{v}^{i}\right) = \mathbf{0}$$

Equation of Continuity

• Neglect  $\eta, \zeta$  and take non-relativistic limit:

$$\left(g_{\nu}^{i}-u_{\nu}u^{i}\right)\partial_{\mu}T^{\mu\nu}=0\rightarrow\rho\left(\partial_{t}v^{i}+v^{j}\partial_{j}v^{i}\right)=-\partial^{i}P$$
  
Euler Equation

• Include  $\eta, \zeta$ :

$$\left(\boldsymbol{g}_{\nu}^{i}-\boldsymbol{u}_{\nu}\boldsymbol{u}^{i}\right)\partial_{\mu}\boldsymbol{T}^{\mu\nu}=\boldsymbol{0}
ightarrow$$
 Navier – Stokes Equation

#### Non-relativistic Ideal Fluid Dynamics

$$\partial_t \mathbf{v}^i + \mathbf{v}^m \partial_m \mathbf{v}^i = -\frac{1}{\rho} \partial_j \delta^{ij} \mathbf{p}$$

[L. Euler, 1755]

- Non-linear
- Non-dissipative: "Ideal Fluid Dynamics"

# Non-linear & Non-dissipative: Turbulence



# Non-linear & Dissipative: Laminar





Non-linear & Dissipative: Laminar

# Viscosity dampens turbulent instability!

# Strongly Coupled Systems

- Consider dimensionless ratio η/s where s is entropy density.
- Generally

$$\frac{\eta}{{\pmb s}} \propto \lambda$$

where  $\lambda$  is mean free path

► Strong interactions mean small  $\lambda$ . Small  $\eta/s$  means good fluid behavior

Strongly Coupled Systems Are Good Fluids!

# Heavy-Ion Colliders: Relativistic Heavy-Ion Collider



















# Experimental Observables



# dN/dp/dø Ø

# Experimental Observables

#### **Experimental Observables**

For ultrarelativistic heavy-ion collisions,

$$\frac{dN}{dp_{\perp}d\phi dy} = \langle \frac{dN}{dp_{\perp}d\phi dy} \rangle_{\phi} \left(1 + 2\nu_2(p_{\perp})\cos(2\phi) + \ldots\right)$$

- Radial flow:  $\langle \frac{dN}{dp_{\perp}dy} \rangle_{\phi}$
- Elliptic flow:  $v_2(p_{\perp})$

#### **Experimental Data**

Hadron Spectra at RHIC  $\sqrt{s} = 200 \text{ GeV}$ 



[PHENIX, 2004]

#### Strongly Coupled Systems — Example RHIC



[Luzum & Romatschke, 2008]

#### Viscosity over Entropy Density: Various Results



[Kovtun, Son, Starinets 2005]

#### Strongly Coupled Systems — Example RHIC

- (One) Current Problem: "Freeze-Out"
- Freeze-Out: converting fluid to particle

$$T_{\rm fluid}^{\mu\nu} = T_{\rm particle}^{\mu\nu} = \int \frac{d^3p}{(2\pi)^3} \sum_i \delta(p^2 - m_i^2) p^{\mu} p^{\nu} f(x, p)$$

But: where to make transition? How treat multi-particle freeze-out?

Experimental setup:

- ► 50-50 mixture of two lowest hyperfine states of <sup>6</sup>Li
- magnetically tuned to a broad Feshbach resonance
- cooled by evaporation in an optical trap
- initial energy per particle *E*<sub>0</sub> measured from trapped cloud profile

#### Attempted fluid description — Ideal Fluids

Ideal Fluids

Continuity Equation

$$\partial_t \rho + \partial_i \left( \rho \mathbf{v}^i \right) = \mathbf{0}$$

• Euler Equation in Trapping potential V(x)

$$\rho\left(\partial_t \vec{v} + \vec{v} \cdot \vec{\partial} \vec{v}\right) = -\vec{\partial} P - \rho \vec{\partial} V$$

• Relation of mass and number density:  $\rho = m \times n$ .

#### Attempted fluid description — Ideal Fluids Initial State: hydrostatics

$$\vec{\partial} P = mn \vec{\partial} V$$

Thermodynamics:

$$\epsilon + P = T s + \mu n$$
,  $d\epsilon = T ds + \mu dn$ .

implies  $dP = nd\mu$ 

$$\mu = \mu_0 + mV(x)$$

• Typically V(x) harmonic trap potential

$$V(x) = \frac{1}{2}\omega_i^2 x_i^2$$

SO

$$\mu = \mu_0 \left( 1 - \frac{x_i^2}{R_i^2} \right) , \quad R_i^2 = \frac{2\mu_0}{m\omega_i^2}$$

#### Equation of State

- ► 4 hydrodynamic equations (Continuity+Euler), 5 variables: n, v, P. Need equation of state to close system
- EoS in the unitarity limit:

$${\cal P}(\mu,T)=\mu^{5/2}m^{3/2}g\left(rac{T}{\mu}
ight)$$

where  $g(z) \rightarrow \text{const.}$  for low temperatures  $T \ll T_F = \frac{(3\pi^2 n)^{2/3}}{2m}$ . Then  $n = \frac{\partial P}{\partial u} \propto \mu^{3/2}$ 

Note that we don't want to enter superfluidity, which happens at

$$T_c \lesssim 0.15 T_F$$



[Cao et al., Science 2011]



[T. Schäfer, 2010]



[Cao et al., Science 2011]



[Kinast, Turlapov, Thomas 2004]

Strongly Coupled Systems — Example Unitary Fermi gas



[Cao et al., Science 2011]

#### Same tool – same problem

- Hydrodynamics not valid in dilute corona
- Qualitatively similar to problem encountered in HIC
- Same problem also in other fields: neutrino sphere in supernovae, surface of last scattering in cosmology,...

# Hydrodynamics for strongly coupled systems

- Strongly coupled systems are good fluids
- Fluid dynamics can be used to extract material properties of strongly coupled systems
- Several examples: QGP, Cold Atoms, Graphene, neutron stars(?)

#### Challenges

- Use different systems to improve fluid dynamics modelling (for example freeze-out)
- Look for universal features
- Learn something about turbulence?

#### Conclusions

- Hydrodynamics is good tool for strongly coupled systems
   provided one is not interested in fast, small details
- Research on hydrodynamics continues Lots of things to be done!