







### Duality sum rules in forward Compton scattering and the proton radius puzzle



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MG, Hobbs, Londergan, Szczepaniak, Phys.Rev. C84 (2011) 065202 MG, Llanes-Estrada, Szczepaniak, Phys.Rev. A87 (2013) 052501, [arXiv:1302.2807] Carlson, MG, Vanderhaeghen, Phys.Rev. A89 (2014) 022504, [arXiv:1311.6512] MG, Phys.Rev. C90 (2014) 052201, [arXiv:1406.1612] MG, Phys.Rev.Lett 115 (2015) 222503, [arXiv:1508.02509]

#### As a motivation

# Proton Radius Puzzle: the Status



### Proton radius puzzle



# Elastic Electron Scattering



Measure cross section down to low  $Q^2$ 

$$\frac{d\sigma^{exp}}{d\Omega} / \left(\frac{d\sigma}{d\Omega}\right)_{Mott} \Big|_{Q^2 \to 0} = 1 + Q^2 \left[\frac{\mu_p^2 - 1}{4M^2} - \frac{1}{3}R_{Ch}^2\right] + \dots$$

The radius is defined as the slope of the FF at origin, data are at finite  $Q^2$ : extrapolation is unavoidable

How low in Q<sup>2</sup> should/can one go? up to now  $Q_{min}^2 = 4 \times 10^{-3} \text{ GeV}^2$ 

1% uncertainty in  $R_{Ch}$  – measure 1 to few x 10<sup>-4</sup> precision!

A1 @ MAMI R<sub>Ch</sub> = 0.879(8) Bernauer et al., '10



- Individual data points per cent level accuracy;
- Need large angle coverage to extract the radius to 1%
- Large statistics serves as a lever arm for extracting "1"to 0.05% precision;
- Higher Q<sup>2</sup> data influence the extracted radius

• The lower in  $Q^2$  one goes, the lesser are higher order terms important – plans with ISR @ Mainz, PRad @ JLab,  $Q^2 \ge 10^{-4} \text{ GeV}^2$ 

- Bernauer et al.: used full statistics (low and moderate  $Q^2$ ) studied systematics due to different fit functions (polynomial, splines, dipole, double dipole etc.)  $R_E^P = 0.879(8)$  fm  $\chi^2$  close to 1 with 1400 d.o.f.

- Lorenz '12,13: Dispersion relation fit  $G_{E,M}(Q^2) = \int_{4m^2}^{\infty} \frac{dt \, \rho_{E,M}(t)}{t+Q^2}$ 

Model of the spectral function:  $2\pi$  continuum + VDM + QCD asymptotics Radius mainly sensitive to the lowest states ( $2\pi$ ,  $3\pi$ ) which are taken as exact -> fit function might not be flexible enough,  $\chi^2 > 1.1$ Consistent with previous DR fits (Höhler '76, Mergell '96, ...)

 $R_{E}^{P} = 0.84(1) \text{ fm}$ 

- Hill, Paz '10: Conformal mapping + Fourier series for the spectral fn.  $R_E^P = 0.87(2)$  fm Data tend to larger radii; Need extra input to get smaller radii

No extrapolation problem in atoms; typical momentum transfer in H-atom: keV<sup>2</sup> in e-H, MeV<sup>2</sup> µ-H



Electrons occupy stationary orbits Energy levels  $E_{NL}$ 

Principal (energy) Q.N.: N=1,2,3...; Orbital momentum Q.N.: L=S,P,D...;

If only one photon were exchanged:  $E_{2S} = E_{2P}$ 





nS-nP splitting (Lamb shift) – authentic prediction of SM (QED) Precise calculations of QED corrections: p.p.m. level precision

The proton is not a point-like charge – has a finite size
 Lamb shift is sensitive to the proton radius

$$\Delta E_{nP-nS} = \Delta E_{nP-nS}^{QED} - \frac{2(Z\alpha)^4}{3n^3} m_r^3 R_E^2 + \mathcal{O}(\alpha_{em}^5)$$

few p.p.m. correction
exceeds the QED precision
can be extracted

 $E_{2S} - E_{2P} = 33.7808(1) \,\mu \text{eV} + 0.0008 R_E^{p\,2} \,\mu \text{eV}$ QED Finite Size

CODATA  $R_{Ch} = 0.8779(94) \text{ fm}$ e-scattering  $R_{Ch} = 0.879(8) \text{ fm}$ Combined  $R_{Ch} = 0.8775(51) \text{ fm}$  $\mu$ H data @ PSI

 $R_E^p = 0.84087(39) \,\mathrm{fm}$ 



Pohl et al [CREMA Coll.] '10, Antognini et al. '13

4% discrepancy for  $R_{Ch}$  (0.6% precision from e-p) - 7 $\sigma$  away!

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 $R^p_E = 0.84087(39) \, \mathrm{fm}$  Pohl et al [CREMA Coll.] '10, Antognini et al. '13 4% discrepancy for R<sub>Ch</sub> (0.6% precision from e-p) – 7 $\sigma$  away!

### RE<sup>P</sup> from e-H

Almost all individual e-H points are within  $1.5\sigma$  from the muonic point BUT they all lie systematically at larger radii – correlated systematics? All QED corrections have been studied up to  $\alpha^6$  – under control Electron scattering is the most precise <u>single</u> measurement and is in nice agreement with the statistical average of the e-H data.

Most of the measurements are old – may be a good idea to remeasure New experiments with projected 1% radius extraction – under way: 2S-2P measurement – York U. (Canada); 2S-4S measurement – MPI Garching; 1S-3S measurement – Laboratoire Kastler Brossel (Paris);

## What's special about $\mu$ -H?

QED: the only difference is the mass

Hydrogen atom



Bohr radius  $R_B \sim \frac{1}{\alpha m_r}$   $m_{\mu} \approx 200 \, m_e$ 

muonic Hydrogen



Fine structure constant Reduced lepton-proton mass  $m_r = \frac{mM}{m+M}$ 

 $\alpha \approx 1/137$ 

Finite size Lamb shift:

 $\Delta E_{2P-2S}^{R_E^p} \propto \alpha^4 m_r^3$ 

 $\Delta E_{2P-2S}^{eH} = -8.1 \times 10^{-7} R_E^2 \text{ meV} \qquad \Delta E_{2P-2S}^{\mu H} = -5.2275(10) R_E^2 \text{ meV}$  $\mu$ H unstable ( $\tau_{2S} \sim \mu$ s) – 7 o.o.m. still make it 10 times more precise

### $R_E^P$ from $\mu$ -H

# Using the proton radius from eH and scattering, expect $\begin{bmatrix} \Delta E_{2P-2S}^{\text{Measured}} - \Delta E_{2P-2S}^{QED} \end{bmatrix}^{\text{Expected}} \approx -4.0 \text{ meV}$ Observed splitting – off by 8%, radius off by 4%

 $\left[\Delta E_{2P-2S}^{\text{Measured}} - \Delta E_{2P-2S}^{QED}\right]^{\text{Measured}} \approx -3.7 \,\text{meV}$ 

What if the  $\mu$ H experiment is wrong? Exp. precision:  $\mu$ eV, much smaller than missing 300  $\mu$ eV; Pohl et al. and Antognini et al. measured  $2P_{1/2} - 2S$  and  $2P_{3/2} - 2S$  transitions, found consistency; No other facility able to redo the  $\mu$ H experiment exists at the moment.



Correction 0.03 meV - 10 times smaller than the discrepancy

### Proton Radius Puzzle: New Physics?





Stringent constraints from  $(g-2)_e$ : substantial  $\mu$ -e non-universality



### Proton Radius Puzzle: New Physics?

#### K-decay constraints

but fine tuned to evade the q-2 constraints



- Solid line is sum of scalar and pseudoscalar couplings.
- Lower mass or higher mass o.k., but 90–200 MeV excluded.
- Same for polar and axial vectors.
- Solid is one particle with both V and A couplings.
- Dashed line is two particles, one polar and one axial vector.
- Lower masses excluded, 160 MeV for PV case, 210 for other case.

Conclusion: BSM explanation possible, requires lepton non-universality,

#### Carlson, Rislow, '12

### Further hadronic effects?

#### Hadronic correction at $(Z\alpha)^5$ – included partially!

Soft Coulomb: Schrödinger WF



Hard box: only part of it included (3<sup>rd</sup> Zemach m.)



#### Do the full calculation



# Blob: forward virtual Compton tensor $T_{\mu\nu} = \frac{i}{8\pi M} \int d^4x e^{iqx} \langle p|T j_{\mu}(x)j_{\nu}(0)|p\rangle$

$$M_{2\gamma} = e^4 \int \frac{d^4q}{(2\pi)^4} \frac{1}{q^4} \bar{u}(k) \left[ \gamma^{\nu} \frac{1}{\not{k} - \not{q} - m_l + i\epsilon} \gamma^{\mu} + \gamma^{\mu} \frac{1}{\not{k} + \not{q} - m_l + i\epsilon} \gamma^{\nu} \right] u(k) T_{\mu\nu}$$

### Polarizability Correction from DR

$$T_{\mu\nu} = \frac{i}{8\pi M} \int d^4x e^{iqx} \langle p|T j_{\mu}(x)j_{\nu}(0)|p\rangle$$

T-ordered non-local product of two vector currents - complicated!

Gauge, Lorentz inv.  $T^{\mu\nu} = \left(-g^{\mu\nu} + \frac{q^{\mu}q^{\nu}}{q^2}\right)T_1(\nu, Q^2) + \frac{\hat{p}^{\mu}\hat{p}^{\nu}}{M^2}T_2(\nu, Q^2)$ 

#### (nP - nS) splitting

$$\Delta E = -\frac{\alpha^2}{2\pi m_l M_d} \phi_n^2(0) \int d^4q \frac{(q^2 + 2\nu^2)T_1(\nu, q^2) - (q^2 - \nu^2)T_2(\nu, q^2)}{q^4[(q^2/2m_l)^2 - \nu^2]}$$

# Polarizability Correction from DR

Optical theorem: absorptive part of  $T_{1,2}$  related to data



Form factors

Unpolarized structure functions F<sub>1,2</sub>

Dispersion relations (subtracted for  $T_1$ )

 $\operatorname{Re} T_{1}(\nu, Q^{2}) = T_{1}(0, Q^{2}) + \frac{\nu^{2}}{2\pi M} \mathcal{P} \int_{0}^{\infty} d\nu' \frac{F_{1}(\nu', Q^{2})}{\nu'(\nu'^{2} - \nu^{2})}$  $\operatorname{Re} T_{2}(\nu, Q^{2}) = \frac{1}{2\pi} \mathcal{P} \int_{0}^{\infty} d\nu' \frac{F_{2}(\nu', Q^{2})}{(\nu'^{2} - \nu^{2})}$ 

### Polarizability Correction

#### Dispersion Relation + Data



# Lamb shift is obtained as $\Delta E \sim \alpha_{em}^5 \int_0^\infty dQ^2 \int_0^\infty d\nu \left\{ A(\nu, Q^2) F_1 + B(\nu, Q^2) F_2 \right\}$

Good quality data (e.g., JLab) on  $F_{1,2}$  O<  $Q^2$ < 3 GeV<sup>2</sup>, W< 4 GeV

### Polarizability Correction



Subtraction function related to proton's magnetic polarizability  $\beta_M$ Low-Energy Theorem: T<sub>1</sub>(0, Q<sup>2</sup>) = (Q<sup>2</sup>/e<sup>2</sup>)  $\beta_M$ 

Lamb shift is obtained as  $\Delta E^{Sub} \sim \alpha_{em}^5 \int_0^{\infty} dQ^2 C(Q^2) \beta_M F_\beta(Q^2)$ 

#### Subtraction Constant



PDG 2012  $\alpha_E = 11.2(0.4) \times 10^{-4} \text{fm}^3$  $\beta_M = 2.5(0.4) \times 10^{-4} \text{fm}^3$ 



### Total polarizability correction

Different approaches to estimate  $F_{\beta}(Q^2)$ 

Dipole (like FF): Pachucki, 1996 Pion loops: Vanderhaeghen & Carlson, 2011 HBChPT + dipole: Birse & McGovern, 2012 BChPT: Alarcón, Pascalutsa,Lenski 2014 Finite Energy Sum Rule: MG, Llanes-Estrada, Szczepaniak, 2013

Hadronic structure corrections to proton radius puzzle are constrained

$$\Delta E_{2P-2S} = -40 \pm 5 \,\mu\text{eV}$$

$$\downarrow$$

$$\Delta E_{\text{Missing}} \approx -300 \,\mu\text{eV}$$

All known constraints built in!

### Exotic Hadronic Contributions?

#### Reasonable hadronic models



#### To get -300 µeV Lamb shift: need something like this



### **Exotic Hadronic Contributions?**

Cottingham formula (p-n n

 $M_p - M_n = \frac{\alpha}{2M(2\pi)^3} \int \frac{d^4q}{q^2} \left[ I^p_{\mu}(\nu, q) - I^{\mu}_{\mu}(\nu, q) \right]$ 



Subtraction function contribution  $[M_p - M_n]^{Subt}$ 

$$= -\frac{\beta_{M}^{p} - \beta_{M}^{n}}{(8\pi)^{2}M} \int_{0}^{\Lambda^{2}} dQ^{2}Q^{2}F_{\beta}(Q^{2})$$

If the proton radius puzzle is due to subtraction contribution  $\delta M_{em}^p \sim 600 \, MeV$ 

Could be purely isoscalar but... VERY unnatural! Should be seen in Deuteron (I=0)



#### Muonic deuterium

One further piece of information available – isotope shift: simultaneous 1S-2S splitting measurement in eH and eD  $R_d^2 - R_p^2 = 3.82007(65) \,\mathrm{fm}^2$  $R_d^2 - R_p^2$  from  $\mu$ H,  $\mu$ D @ PSI – in agreement (preliminary) Exotic hadronic contributions excluded by this finding

Extraction from  $\mu$ D relies on nuclear structure-dependent polarizability correction.

Nuclear models vs dispersion relations:

 $\Delta E_{2S}^{Nucl.} = -1.68(16) \,\mathrm{meV}$ 

Leidemann, '90; Pachucki '13; Ji et al, '14; Friar, '14;  $\Delta E_{2S}^{DR} = -1.75(74) \text{ meV}$ Carlson, MG, Vanderhaeghen '14 A simple ansatz for F<sub>β</sub>(Q<sup>2</sup>) used

### Lacking Input to DR for $\mu D$

 $\Delta E \sim \alpha_{em}^5 \int_0^\infty dQ^2 \int_0^\infty d\nu \left\{ A(\nu, Q^2) F_1 + B(\nu, Q^2) F_2 \right\}$ All kinematics contribute to the dispersive integral; Not all of them are equally important

The bulk of the correction – quasi elastic data from  $v \approx 6-10$  MeV and Q<sup>2</sup> < 0.005 GeV<sup>2</sup> – just below the kinematics of available QE data

New D(e,e')pn data down to  $Q^2 = 0.002 \text{ GeV}^2 \text{ A1}@\text{MAMI}$ taken and under analysis; 2% measurement will reduce the uncertainty by a factor 2-4

Once the data are more precise: the model for  $F_{\beta}(Q^2)$  will become the main limitation of the calculation

### Subtraction function from finite energy sum rule

 $\operatorname{Re} T_1(\nu, Q^2) = T_1(0, Q^2) + \frac{\nu^2}{2\pi M} \mathcal{P} \int_0^\infty d\nu' \frac{F_1(\nu', Q^2)}{\nu'(\nu'^2 - \nu^2)}$ 

#### FESR (real photons)

Nuclear photoabsorption: from  $v_{thr}$  = few MeV to  $v_{max}$  = few tens MeV; "nothing" above that until  $v_{\pi}$  = 150 MeV; Scale separation:  $v_{max} \ll v_{\infty} \ll v_{\pi}$ 



Evaluate the DR at  $\nu = \nu_{\infty}$ Re  $T_1(\nu_{\infty}, 0) = \operatorname{Re} T_1(0, 0) + \frac{\nu_{\infty}^2}{2\pi M} \mathcal{P} \int_{\nu_{thr}}^{\infty} \frac{d\nu}{\nu(\nu^2 - \nu_{\infty}^2)} F_1(\nu, 0)$ Employ duality LEX at  $\nu$ =0: nuclear Thomson term  $\operatorname{Re} T_1(0, 0) = -\frac{Z^2}{4\pi M}$ LEX at  $\nu$ = $\nu_{\infty}$ : nucleon Thomson terms + polarizabilities  $\operatorname{Re} T_1(\nu_{\infty}, 0) = -\frac{Z}{4\pi M_p} + \frac{\nu_{\infty}^2}{e^2} (Z(\alpha^p + \beta^p) + N(\alpha^n + \beta^n))$ 

 $\frac{\nu_{\infty}^2}{2\pi M} \mathcal{P} \int \frac{d\nu}{\nu(\nu^2 - \nu_{\infty}^2)} F_1(\nu, 0) \approx$ 

Work out the integral

$$= -\frac{1}{2\pi M} \int_{\nu_{thr}}^{\nu_{max}} \frac{d\nu}{\nu} F_1(\nu, 0)$$

$$= \frac{\nu_{\infty}^2}{2\pi M} \mathcal{P} \int_{\nu_{max}}^{\nu_{\pi}} \frac{d\nu}{\nu(\nu^2 - \nu_{\infty}^2)} F_1(\nu, 0)$$

$$= \frac{\nu_{\infty}^2}{2\pi M} \int_{\nu_{max}}^{\infty} \frac{d\nu}{\nu^3} F_1(\nu, 0)$$

Balance of coeffs. at  $(v_{\infty})^2$ :

L.H.S.  $(\nu_{\infty}^{2}/e^{2})[Z(\alpha^{p}+\beta^{p})+N(\alpha^{n}+\beta^{n})]$ 

Baldin sum rule for nucleons:

$$\alpha_E^{p,n} + \beta_M^{p,n} = \frac{2\alpha}{M} \int\limits_{\mathcal{U}}^{\infty} \frac{d\nu}{\nu^3} F_1^{p,n}(\nu,0)$$

 $\mathcal{P}\int_{\nu_{thr}}^{\infty} \frac{d\nu}{\nu(\nu^2 - \nu_{\infty}^2)} F_1(\nu, 0) + \int_{\nu_{\pi}}^{\infty} \frac{d\nu}{\nu^3} \left[ F_1(\nu, 0) - (Z + N)(ZF_1^p(\nu, 0) + NF_1^n(\nu, 0)) \right] \approx 0$ 

Non-interacting nucleons in the nucleus

#### Coeffs. at $(\nu_{\infty})^{0}$ : Bethe-Levinger photonuclear sum rule

$$-\frac{Z}{4\pi M_p} = -\frac{Z^2}{4\pi M} - \frac{1}{2\pi M} \int_{\nu_{thr}}^{\nu_{max}} \frac{d\nu}{\nu} F_1(\nu, 0)$$

$$2\int_{\nu_{thr}}^{\nu_{max}} \frac{d\nu}{\nu} F_1(\nu, 0) = ZN$$

Integrated nuclear photoabsorption cross section is given by the number of "elementary" scatterers – nucleons

Thomas – Reiche – Kuhn sum rule in QM: integrated oscillator strength ~ number of oscillators

#### Bethe-Levinger SR: works to 10-20%

#### 740

B. L. Berman and S. C. Fultz: Measurements of the giant dipole resonance

|                   |                         | 100 C                               |                             | σ.   |   |                                    |
|-------------------|-------------------------|-------------------------------------|-----------------------------|--|---|------------------------------------|
|                   |                         | $\sigma_{\rm int}(\gamma,{ m tot})$ |                             |  | $\sigma_{\rm int}[(\gamma,2n)+(\gamma,3n)]$ |                                    |
| Nucleus           | $E_{\gamma \max}$ (MeV) | 60NZ/A                              | $\sigma_{-1} A^{-4/3}$ (mb) | $\begin{array}{c} 0.00225 \ A^{5/3} \\ (\text{mb-MeV}^{-1}) \end{array}$ | $\sigma_{\rm int}(\gamma,{ m tot})$         | Reference                          |
| <sup>91</sup> Zr  | 30.0                    | 0.820                               | 0.160                       | 0.98   | 0.181                                       | Berman et al., 1967                |
| <sup>92</sup> Zr  | 27.8                    | 0.804                               | 0.154                       | 0.93   | 0.414                                       | Berman et al., 1967                |
| <sup>93</sup> Nb  | 24.3                    | 0.967                               | 0.186                       | 1.12   | 0.209                                       | Leprêtre et al., 1971              |
| <sup>94</sup> Zr  | 31.1                    | 0.813                               | 0.160                       | 1.01   | 0.547                                       | Berman et al., 1967                |
| <sup>107</sup> Ag | 29.5                    | 0.858                               | 0.155                       | 0.89   | 0.194                                       | Berman et al., 1969a               |
| 115In             | 31.1                    | 1.111                               | 0.202                       | 1.17   | 0.278                                       | Fultz et al., 1969                 |
| 116Sn             | 29.6                    | 0.978                               | 0.175                       | 0.99   | 0.248                                       | Fultz et al., 1969                 |
| 117Sn             | 31.1                    | 1.102                               | 0.199                       | 1.16   | 0.271                                       | Fultz et al., 1969                 |
| 118Sn             | 30.8                    | 1.072                               | 0.190                       | 1.07   | 0.297                                       | Fultz et al., 1969                 |
| 119Sn             | 31.1                    | 1.145                               | 0.202                       | 1.17   | 0.334                                       | Fultz et al., 1969                 |
| <sup>120</sup> Sn | 29.9                    | 1.185                               | 0.209                       | 1.19   | 0.330                                       | Fultz et al., 1969                 |
| <sup>124</sup> Sn | 31.1                    | 1.123                               | 0.200                       | 1.16   | 0.361                                       | Fultz et al., 1969                 |
| $^{127}I$         | 29.5                    | 0.933                               | 0.164                       | 0.93   | 0.256                                       | Bramblett et al., 1966b            |
|                   | 24.9                    | 1.074                               | 0.201                       | 1.18   | 0.196                                       | Bergère et al., 1969               |
| <sup>133</sup> Cs | 29.5                    | 1.026                               | 0.182                       | 1.04   | 0.257                                       | Berman et al., 1969a               |
| <sup>138</sup> Ba | 27.1                    | 1.022                               | 0.183                       | 1.05   | 0.242                                       | Berman et al., 1970c               |
| <sup>139</sup> La | 24.3                    | 0.980                               | 0.177                       | 1.02   | 0.147                                       | Beil et al., 71                    |
| <sup>141</sup> Pr | 29.8                    | 1.001                               | 0.175                       | 0.97   | 0.167                                       | Bramblett et al., 1966b            |
|                   | 16.9                    | 0.691                               | 0.138                       | 0.85   |   | Beil et al., 1971                  |
|                   | 18.1                    | 0.678ª                              | 0.128ª                      | 0.75ª  |   | Young, 1972                        |
| $^{142}Nd$        | 20.2                    | 0.901                               | 0.170                       | 1.00   | 0.024                                       | Carlos et al., 1971                |
| <sup>143</sup> Nd | 19.8                    | 0.910                               | 0.176                       | 1.08   | 0.094                                       | Carlos et al., 1971                |
| 144Nd             | 20.2                    | 0.896                               | 0.170                       | 1.01   | 0.299                                       | Carlos et al., 1971                |
| <sup>145</sup> Nd | 20.2                    | 0.965                               | 0.193                       | 1.26   | 0.323                                       | Carlos et al., 1971                |
| 146Nd             | 20.2                    | 0.905                               | 0.173                       | 1.05   | 0.347                                       | Carlos et al., 1971                |
| <sup>148</sup> Nd | 18.8                    | 0.795                               | 0.155                       | 0.97   | 0.491                                       | Carlos et al., 1971                |
| <sup>150</sup> Nd | 20.2                    | 0.931                               | 0.178                       | 1.09   | 0.416                                       | Carlos et al., 1971                |
| <sup>153</sup> Eu | 28.9                    | 1.022                               | 0.181                       | 1.03   | 0.311                                       | Berman et al., 1969b               |
| 199 L.P           | 28.0                    | 0.997                               | 0.175                       | 1.00   | 0.386                                       | Bramblett et al., 1964             |
|                   | 27.4                    | 1.109                               | 0.198                       | 1.15   | 0.243                                       | Bergère et al., 1968               |
| 160Gd             | 29.5                    | 1.099                               | 0.195                       | 1.14   | 0.448                                       | Berman <i>et al.</i> , 1969b       |
| <sup>165</sup> Ho | 28.9                    | 1.057                               | 0.183                       | 1.04   | 0.312                                       | Berman et al., 1969b               |
|                   | 26.8                    | 1.202                               | 0.215                       | 1.24   | 0.272                                       | Bergère et al., 1968               |
| 175Lu             | 23.0                    | 0.990                               | 0.177                       | 1.02   | 0.253                                       | Bergère et al., 1969               |
| <sup>181</sup> Ta | 24.6                    | 0.835                               | 0.146                       | 0.82   | 0.404                                       | Bramblett et al., 1963             |
| -                 | 25.2                    | 1.142                               | 0.201                       | 1.14   | 0.269                                       | Bergère et al., 1968               |
| 186W              | 28.6                    | 1.123                               | 0.191                       | 1.06   | 0.449                                       | Berman et al. 1969b                |
| <sup>197</sup> Au | 24.7                    | 1.045                               | 0.179                       | 0.98   | 0.262                                       | Fultz et al., 1962h                |
|                   | 21:7                    | 1,080                               | 0 190                       | 1 06   | 0.156                                       | Vevssière $\rho t \sigma l = 1070$ |
| 206Pb             | 26.4                    | 0.982                               | 0 167                       | 0.03   | 0 183                                       | Horvey et al. $1064$               |
| <b>.</b> <i>D</i> | 20.1                    | 0.304                               | 0.107                       | 0.90   | 0.105                                       | 11aivey 61 ul., 1904               |

TABLE III. Quantities derived directly from the data—all nuclei.

Rev. Mod. Phys., Vol. 47, No. 3, July 1975

#### Include hadronic photoabsorption

Complication: c.s. increases at high energies

$$F_1(\nu \ge 2 \operatorname{GeV}, 0) \to F_1^R(\nu, 0) = C_M \left(\frac{\nu}{\nu_0}\right)^{\alpha_M} + C_P \left(\frac{\nu}{\nu_0}\right)^{\alpha_F}$$

 $\nu_0 \approx 1 \,\text{GeV}, \quad \alpha_M \approx 0.5, \quad \alpha_P \approx 1.09$ 



Build a Regge-behaved analytic function  $\operatorname{Re} T_1^R(\nu, 0) = 0 + \frac{\nu^2}{2\pi M} \mathcal{P} \int_0^\infty \frac{d\nu'}{\nu'(\nu'^2 - \nu^2)} F_1^R(\nu', 0)$ 

Subtract Regge behavior: the integral runs up to finite energy N

$$\operatorname{Re}\left[T_{1}(\nu,0) - T_{1}^{R}(\nu,0)\right] = -\frac{Z^{2}}{4\pi M} + \frac{\nu^{2}}{2\pi M} \mathcal{P}\int_{\nu_{thr}}^{N} \frac{d\nu' [F_{1}(\nu',0) - F_{1}^{R}(\nu',0)]}{\nu'(\nu'^{2} - \nu^{2})}$$

The remaining amplitude – at most constant asymptotically The asymptotic constant – (hypothetical) J=0 fixed pole  $C_{\infty} = \operatorname{Re}\left[T_{1}(\nu, 0) - T_{1}^{R}(\nu, 0)\right]_{\nu \to \infty}$ 

Analyticity: the J=0 pole is not a free constant

$$C_{\infty} = -\frac{Z^2}{4\pi M} - \frac{1}{2\pi M} \int_{\nu_{thr}}^{N} \frac{d\nu}{\nu} F_1(\nu, 0) + \frac{1}{2\pi M} \left[ \frac{C_M}{\alpha_M} \left( \frac{N}{\nu_0} \right)^{\alpha_M} + \frac{C_P}{\alpha_P} \left( \frac{N}{\nu_0} \right)^{\alpha_P} \right]$$

Damashek and Gilman, 1969

Exact duality: integrated c.s. = integrated Regge

-> J=0 pole = Thomson term

Deviation of J=0 pole from Thomson term = duality violation



cf. Müller, Polyakov, Semenov 2015

Choose  $v_{\infty} \sim \text{few GeV}(v_{\infty}) \sim N) - \text{sum of CQ Thomson terms}$ 

$$\operatorname{Re}\left[T_1(\nu_{\infty}, 0) - T_1^R(\nu_{\infty}, 0)\right] = -\sum_{q=u, d \in A} \frac{e_q^2}{4\pi M_q} = -\frac{3Z + 2N}{4\pi M_p} \qquad M_q \approx M_p/3$$

"Identify" meson Regge exchange as quark-antiquark exchanges

CQM sum rule 
$$(Z+N)^2 + \frac{ZN}{2} = \int_{\nu_{thr}}^{\nu_{max}} \frac{d\nu}{\nu} F_1(\nu,0) - \frac{C_M}{\alpha_M} \left(\frac{\nu_{max}}{\nu_0}\right)^{\alpha_M}$$

#### Fit of photoabsorption data on a few selected nuclei Resonance + Regge background MG, Hobbs, Londe

MG, Hobbs, Londergan, Szczepaniak 2011







Message: duality sum rules work; Can be used for quantitative study; Precision – can be 10–20% Generalize Bethe-Levinger SR to finite Q<sup>2</sup>:

 $\operatorname{Re} T_1(\nu_{\infty}, Q^2) = \operatorname{Re} T_1(0, Q^2) - \frac{1}{2\pi M} \int \frac{d\nu}{\nu} F_1(\nu, Q^2)$ 

LEX at finite  $Q^2$ : Dirac (or charge) form factor + magnetic pol.

$$T_{1}(0,Q^{2}) = -\frac{Z^{2}}{4\pi M}F_{D}^{2}(Q^{2}) + \frac{Q^{2}}{e^{2}}\beta_{M}F_{\beta}(Q^{2}) \qquad F_{\beta}(0) =$$

$$T_{1}(\nu_{\infty},Q^{2}) = -\frac{Z}{4\pi M_{p}}F_{D}^{p2}(Q^{2}) + Z\frac{Q^{2}}{e^{2}}\beta_{M}^{p}F_{\beta}^{p}(Q^{2}) - \frac{N}{4\pi M_{p}}F_{D}^{n2}(Q^{2}) + N\frac{Q^{2}}{e^{2}}\beta_{M}^{n}F_{\beta}^{n}(Q^{2}) + O(\nu_{\infty}^{2})$$

The new sum rule: the  $Q^2$  slope of the TRK – BL sum rule

$$\beta_M = \frac{2\alpha}{M} \int_{\nu_{thr}}^{\nu_{max}} \frac{d}{d\nu} \left. \frac{d}{dQ^2} F_1(\nu, Q^2) \right|_{Q^2 \to 0}$$
$$- \frac{Z^2 \alpha R_{Ch}^2}{3M} + \frac{Z \alpha R_p^2 + N \alpha R_n^2}{3M_p} + Z \beta_M^p + N \beta_M^n$$

Can test the sum rule:

- fit the electrodisintegration data in the nuclear range;
- compare to the value of the nuclear magnetic pol. (if known)

**Deuteron:**  $\beta_M$  known theoretically

EFT (lowest order):  $\beta_M^d = 0.068 \text{ fm}^2$ Potential models (LO):  $\beta_M^d = 0.068 \text{ fm}^2$ Potential models (NLO):  $\beta_M^d = 0.078 \text{ fm}^2$  Chen et al., 2002 Friar 1997, Khriplovich 1979, ... Friar 1997

Nucleon  $\beta_M$ : known and generally small (2 o.o.m.)PDG 2012 $\beta_M^p = 2.5(0.4) \cdot 10^{-4} \, \text{fm}^3, \quad \beta_M^n = 3.7(2.0) \cdot 10^{-4} \, \text{fm}^3$ ChPT $\beta_M^p = 3.9(0.7) \cdot 10^{-4} \, \text{fm}^3, \quad \beta_M^n = 4.6(2.7) \cdot 10^{-4} \, \text{fm}^3$ 

Hagelstein et al., arXiv:1512.03765

Charge radii known:  $R_d = 2.14 \text{ fm}$ ,  $R_p = 0.840 \text{ fm}$ ,  $R_n^2 = -0.116 \text{ fm}^2$ 

Correction term:

$$-\frac{\alpha R_d^2}{3M} + \frac{\alpha R_p^2 + \alpha R_n^2}{3M_p} + \beta_M^p + \beta_M^n \approx 1 \times 10^{-5} \text{fm}^3$$

Recent deuteron data fit Carlson, MG, Vanderhaeghen, PR A89 (2014) Fit of the form  $F^{QE}(v,Q^2) \cdot f^{QE}(Q^2) + F^{thr}(v,Q^2) \cdot f^{thr}(Q^2)$ 





0.073(5) fm<sup>3</sup> <-> 0.096(16) fm<sup>3</sup>
1.5σ off, but in the ballpark
The problem: need to extrapolate down to Q<sup>2</sup>=0 from finite Q<sup>2</sup>; Nuclear slopes are large

Can impose the value of  $\beta_d$  – new fit

#### Fit done not using the SR



#### New application: He-3 <sup>Carlson, MG, Vanderhaeghen, in progress</sup> Fit of the form $F^{QE}(v,Q^2) \cdot f^{QE}(Q^2) + F^{thr}(v,Q^2) \cdot f^{thr}(Q^2)$



Sum rule prediction for the magnetic polarizability  $\beta_M^{He-3} = [4.20 - 2.44 + 0.67 + 1.24] \cdot 10^{-3} \text{ fm}^3 = 3.9 \cdot 10^{-3} \text{ fm}^3$ Uncertainty? 10% from  $\beta_P$  and  $\beta_n$ ; 10% from the fit; systematics? Further generalization: the full  $Q^2$  dependence of  $\beta(Q^2)$ 

$$\beta_M(Q^2) = \frac{2\alpha}{M} \int_{\nu_{thr}}^{\nu_{max}} \frac{d}{d\nu} \frac{F_1(\nu, Q^2) - F_1(\nu, 0)}{Q^2}$$

Confront to the simple-minded FF-like model of our PR A89 Effect on the Lamb shift calculation Estimate w/o sum rule Estimate with sum rule

 $\Delta E_{2S}^{DR} = -1.75(74) \,\mathrm{meV}$ 

 $\Delta E_{2S}^{SR} = -1.94(74) \,\mathrm{meV}$ 

Uncertainty dominated by the dispersion integral; once more precise data allow to reduce the uncertainty – may lead to a shift in the extracted value of  $R_d$  !

Effect due to different  $R_p \approx 0.38$  meV in  $\mu D$ ; here – 0.19 meV

#### Summary

- Proton radius puzzle inconsistency between the e-scattering and eH on one hand, and  $\mu$ H data on the other hand.
- Each part has subtleties but no clear solution found the puzzle persists
- Scattering experiments: extrapolation issue
- Electronic hydrogen: sensitivity issue
- Muonic hydrogen: no experimental issues found to date further muonic atoms consistent with  $\mu$ H (preliminary)
- SSM explanation possible but requires both lepton non-universality and fine tuning to evade known constraints from other observables

### Proton Radius Puzzle: what's next?

- More precise eH experiments coming (2S-2P, 1S-3S, 2S-4S);
- e-p scattering:  $Q^2$  down to  $2 \times 10^{-4}$  GeV<sup>2</sup> @ Mainz, JLab
- Deuteron radius from e-D scattering: new data at Mainz under analysis Q<sup>2</sup> > 0.002 GeV<sup>2</sup>, radius under 0.25%
- To push Q<sup>2</sup> down and get the radius under 1%: improved radiative corrections (TPE) necessary. Recent works: MG '14; Tomalak, Vanderhaeghen '14, '15(2)
- Study lepton non-universality with μ-p scattering: MUSE @ PSI - elastic μ-p scattering at Q<sup>2</sup> > 0.002 GeV<sup>2</sup> (2017/18);
   γp -> μ<sup>+</sup>μ<sup>-</sup>p/γp -> e<sup>+</sup>e<sup>-</sup>p measurement may be more sensitive Pauk, Vanderhaeghen '15 - proposal under consideration in Mainz

### Proton Radius Puzzle: what's next?

- Further muonic atoms:  $\mu$ D,  $\mu$ He-3,  $\mu$ He-4 data taken at PSI, now analyzed or finalized
- μD more precise DR calculation needed: new QE data on deuteron analyzed at Mainz

   to reduce the uncertainty of dispersion integrals by factor 2-4 sum rule for the nuclear magnetic polarizability derived (MG, '15)
   to reduce model dependence of the subtraction contribution DR treatment of hyperfine splitting in μD underway
   with Carlson and Vanderhaeghen
- $\square$  µHe-3,4 DR analysis underway (with Carlson and Vanderhaeghen) potential model calculation by Bacca and Co arXiv: 1512.05773